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PHYSICS 432/532: Cosmology: HW 3

Solution Key 1. Ryden 6.3 Clarification:

Your textbook uses the phrase

"current proper distance" in place of  
"comoving distance",  $\sim$ . Sol.

Suppose that you are in a spatially at

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universe containing a single component with a unique equation of state parameter,  $w$ . What are the current proper distance  $d_P(t_0)$ , the luminosity distance,  $d_L$  and the angular-diameter ...

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Cosmology – HW 1 Solution Key. 1.  
Ryden 2.2 bf Sol. The density of stars  
is:  $n = 10^9 \text{ Mpc}^{-3} = 3.4 \times 10^{-59} \text{ m}^{-3}$ .

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Introduction To Cosmology Solution  
Manual The book name is only for  
reference purpose.

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cosmology are introduced as needed.

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Unfortunately, the National Bureau of Standards has not gotten around to establishing a standard notation for cosmological equations. It seems that every cosmology book has its own notation; this book is no exception. My main motivation was to make the notation as clear as possible for the



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cosmological novice.

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Introduction to Cosmology -  
University of Arizona

- Barbara Ryden, Introduction to  
Cosmology A clearly written book  
that presents an excellent, gentle

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introduction to the expanding universe, with subsequent chapters on thermal history and structure formation.

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Cosmology - University of Cambridge  
COVERAGE: Lecture Notes 1, 2, and 3;

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Problem Sets 1, 2, and 3; Weinberg, Chapters 1, 2, and 3; Ryden, Chapters 1, 2, and 3. While all of Ryden's Chapter 3 has been assigned, questions on the quiz will be limited to Sections 3.1 (The Way of Newton) and 3.3 (The General Way of Einstein).Section 3.2 (The Special Way

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of Einstein) describes special relativity.

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8.286 Home Page, MIT  
PHY306 Introduction to Cosmology  
Practice Problems NOTE: The  
Friedman equation:  $\dot{a}(t)^2 = 8\pi G \rho - 3c^2 \Lambda$

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$2 \mu E r_0 a(t) + E m_0 a(t) \uparrow - kc^2 R^2_0 +$   
 $3 a(t)$  The fluid equation:  $E' + 3$   
 $a' a (E + P) = 0$  The acceleration  
 equation:  $\ddot{a} a = - \frac{4}{3} G \frac{3c^2}{c^2} (E + 3P)$   
 The Robertson-Walker metric:  $ds^2 =$   
 $- c^2 dt^2 + a(t)^2 \left[ dr^2 + x^2 k_j d\theta^2 + \sin^2 \theta \right.$   
 $d\phi^2 \left. \right]$  where  $x k(r) = R \sin(r/R)$   $k =$   
 $+1$   $r k = 0$   $R \sinh(r/R) \dots$

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exciting update of an award-winning textbook. It is aimed primarily at advanced undergraduate students in physics and astronomy, but is also useful as a supplementary text at higher levels.

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Introduction to Cosmology

Instructor ' s Manual Barbara Ryden

The Ohio State University. 1

Introduction The purpose of this instructor ' s manual is primarily to provide worked so-lutions for the end-of-chapter problems. In writing up the solutions, I didn ' t try to mimic

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the solutions that an A+ student will submit; instead, I tried to go into “ verbose ” mode, giving general solutions ...

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